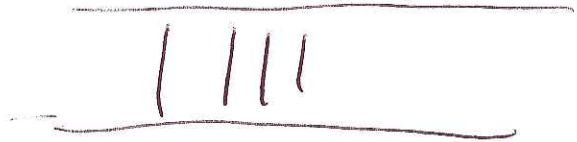


The nuclear atom

1.) History

- a) atomic spectra
- slides

Balmer series



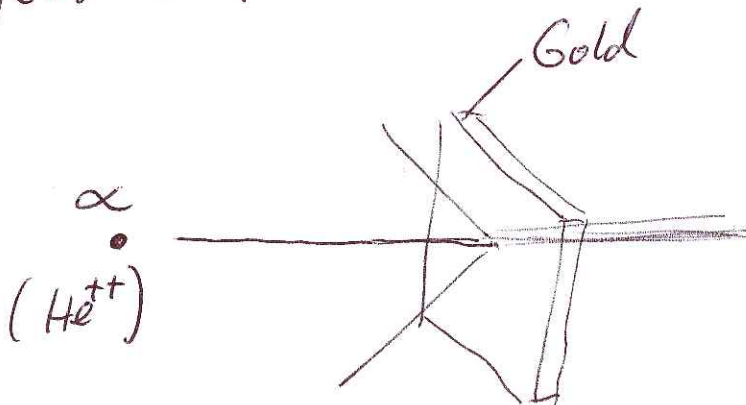
$$\lambda_n = 364.6 \frac{n^2}{n^2 - 4} \text{ nm}$$

General case

$$\frac{1}{\lambda_{nm}} = R \left(\frac{1}{m^2} - \frac{1}{n^2} \right) \quad n > m$$

Rydberg constant

Rutherford

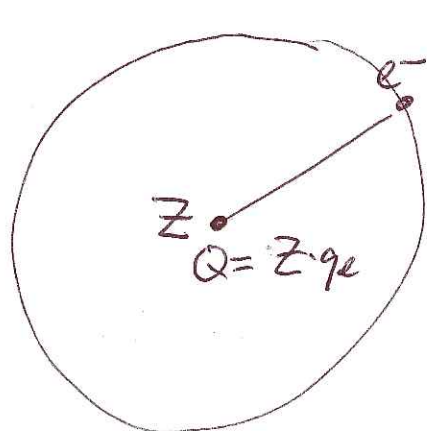


→ Nuclear structure

Data explained by

Coulomb scattering from massive nuclei (Gold)

Bohr model

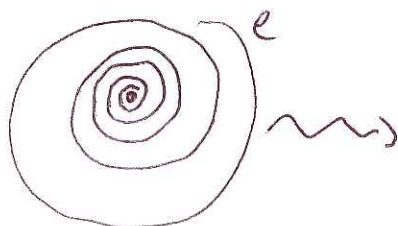


electrons on Keplerian orbit
of central potential

$$F = \frac{kZe \cdot e}{r^2} = \frac{mv^2}{r}$$

*1

Bohr's: Problem: e should radiate



Bohr: $n\lambda = 2\pi r$

$$n\left(\frac{h}{p}\right) = 2\pi r$$

$$n\hbar = \underbrace{pr}_{mvr} = L$$

quantization
of L

$$* m v^2 r = k \cdot Z e^2 \quad (1)$$

$$m^2 v^2 r^2 = n^2 \hbar^2 \quad (2)$$

$$m r = \frac{n^2 \hbar^2}{k Z e^2} \quad \frac{(2)}{(1)}$$

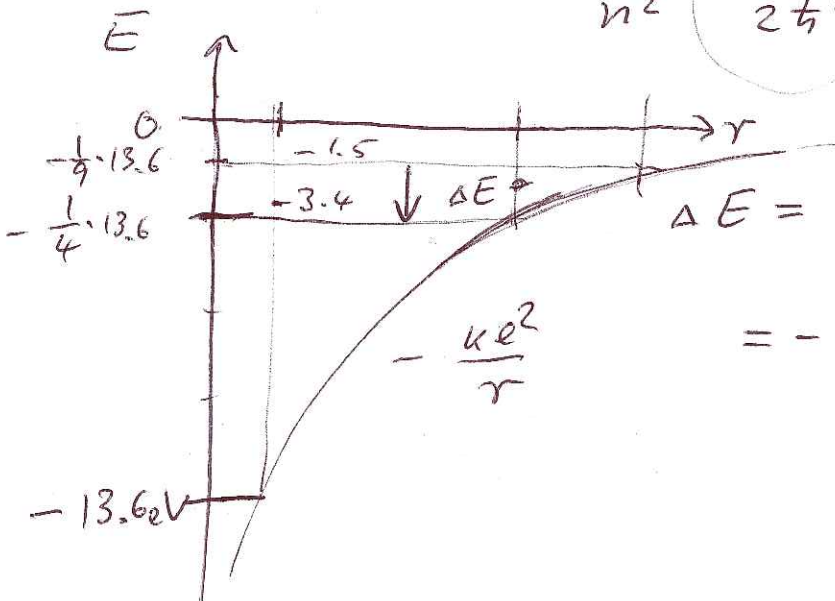
$$\boxed{r = \frac{n^2 \hbar^2}{Z m k e^2}} = n^2 \frac{a_0}{Z}$$

$$a_0 = \frac{\hbar^2}{mke^2} = 0.0529 \text{ nm } (= 0.529 \text{ \AA})$$

Bohr radius

Energy

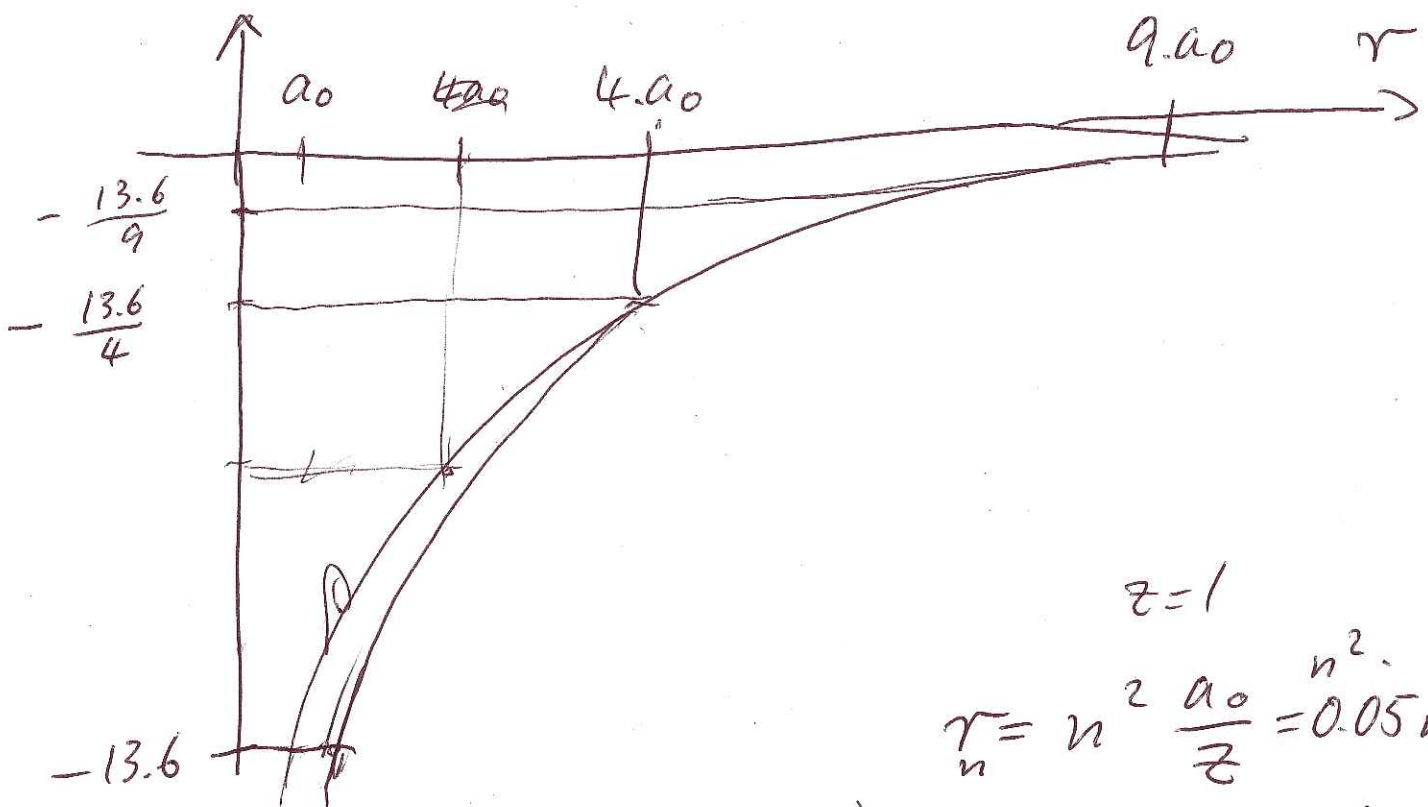
$$\begin{aligned} E &= \frac{p^2}{2m} + V(r) \\ &= \frac{p^2}{2m} - \frac{zke^2}{r} \\ &= \frac{1}{2m} \left(\frac{n\hbar}{r} \right)^2 - \frac{zke^2}{r} \\ &= \frac{1}{2m} \left(\frac{z m k e^2}{n\hbar} \right)^2 - \frac{z k e^2 \cdot z m k e^2}{n^2 \hbar^2} \\ &= - \frac{z^2 m k^2 e^4}{2 n^2 \hbar^2} \\ &= - \frac{z^2}{n^2} \cdot \frac{m k^2 e^4}{2 \hbar^2} \quad 13.6 \text{ eV} \end{aligned}$$



$$\Delta E = E_n - E_m$$

$$= -z^2 \cdot (13.6 \text{ eV}) \left(\frac{1}{n^2} - \frac{1}{m^2} \right)$$

Radii & energy of orbits



$$z=1$$
$$r_n = n^2 \frac{a_0}{z} = 0.05 n^2 \text{ nm}$$

$$E_n = (13.6 \text{ eV}) / n^2$$

